

Skyrme EDF for beyond mean-field calculations

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- Skyrme mean-field and beyond mean-field calculations for masses and excitations
- Standard effective Skyrme interaction: $E = \langle \hat{T} + \hat{V}_2(\rho_0) \rangle$
- Standard effective EDF: $E \neq \langle \hat{T} + \hat{V}_2(\rho_0) \rangle$
- Constraints from beyond mean-field calculations
- Constraints from mean-field calculations
- New Skyrme effective interaction: $E = \langle \hat{T} + \hat{V}_2 + \hat{V}_3 + \hat{V}_4 \rangle$

Microscopic calculation of contributions of *some* correlations on ground states energies

- Mean-field approximation relies on a very simple wave function for the ground state of the system
- Some beyond mean-field effects might be absorbed in an effective way by refitting the coupling constants of the EDF or by using schematic corrections
- Some ground state correlations can be microscopically calculated by going beyond the mean-field approximation by
 - the mechanism of symmetry breaking and restauration
 - the mixing of states along collective coordinates (GCM)

See M. Bender, G.F. Bertsch and P.-H. Heenen, PRC **73**, 034322:

- How large are these correlation energies ?
- How much do they fluctuate ?

■ Effective Skyrme *interaction*

$$\begin{aligned} V_{\text{eff}}(\mathbf{r}) &= t_0 (1 + x_0 \hat{P}^\sigma) \delta(\mathbf{r}) && \text{local} \\ &+ \frac{1}{2} t_1 (1 + x_1 \hat{P}^\sigma) [\mathbf{k}'^2 \delta(\mathbf{r}) + \delta(\mathbf{r}) \mathbf{k}^2] && \text{non local} \\ &+ t_2 (1 + x_2 \hat{P}^\sigma) \mathbf{k}' \cdot \delta(\mathbf{r}) \mathbf{k} && \text{non local} \\ &+ \frac{1}{6} t_3 (1 + x_3 \hat{P}^\sigma) \rho_0^\alpha \delta(\mathbf{r}) && \text{density dep.} \\ &+ i W_0 \hat{\sigma} \cdot [\mathbf{k}' \times \delta(\mathbf{r}) \mathbf{k}] && \text{spin-orbit} \end{aligned}$$

■ Sometimes complemented with a tensor term

■ Possibly complemented with a D-wave term

■ Higher order derivative terms ?

■ Other density dependent terms ?

From effective interaction to EDF (« \hat{T} even » part)

$$\mathcal{H} = \mathcal{K} + \mathcal{H}_0 + \mathcal{H}_3 + \mathcal{H}_{\text{eff}} + \mathcal{H}_{\text{fin}} + \mathcal{H}_{\text{so}} + \mathcal{H}_{\text{sg}} + \mathcal{H}_{\text{coul}}$$

$$\mathcal{H}_0 = \frac{1}{4} t_0 \left[(2 + x_0) \rho_0^2 - (2x_0 + 1) \sum_q \rho_q^2 \right] = \sum_{T=0,1} C_T [\rho_0] \rho_T^2$$

$$\mathcal{H}_3 = \frac{1}{24} t_3 \rho_0^\alpha \left[(2 + x_3) \rho_0^2 - (2x_3 + 1) \sum_q \rho_q^2 \right]$$

$$\mathcal{H}_{\text{eff}} = \frac{1}{8} [t_1(2 + x_1) + t_2(2 + x_2)] \tau_0 \rho_0 = \sum_{T=0,1} C_T^\tau \tau_T \rho_T$$

$$+ \frac{1}{8} [t_2(2x_2 + 1) - t_1(2x_1 + 1)] \sum_q \tau_q \rho_q$$

$$\mathcal{H}_{\text{fin}} = \frac{1}{32} [t_2(2 + x_2) - 3t_1(2 + x_1)] \rho_0 \Delta \rho_0 = \sum_{T=0,1} C_T^{\Delta \rho} \rho_T \Delta \rho_T$$

$$+ \frac{1}{32} [3t_1(2x_1 + 1) + t_2(2x_2 + 1)] \sum_q \rho_q \Delta \rho_q$$

$$\mathcal{H}_{\text{so}} = -\frac{W_0}{2} \left[\rho_0 \nabla \cdot \mathbf{J}_0 + \sum_q \rho_q \nabla \cdot \mathbf{J}_q \right] = \sum_{T=0,1} C_T^{\nabla J} \rho_T \nabla \cdot \mathbf{J}_T$$

$$\mathcal{H}_{\text{sg}} = -\frac{t_1 x_1 + t_2 x_2}{16} \mathbf{J}_0^2 + \frac{t_1 - t_2}{16} \sum_q \mathbf{J}_q^2 = \sum_{T=0,1} C_T^{\mathbf{J}} \mathbf{J}_T^2$$

From effective interaction to EDF (« \hat{T} even » part)

$$\mathcal{H} = \mathcal{K} + \mathcal{H}_0 + \mathcal{H}_3 + \mathcal{H}_{\text{eff}} + \mathcal{H}_{\text{fin}} + \mathcal{H}_{\text{so}} + \mathcal{H}_{\text{sg}} + \mathcal{H}_{\text{coul}}$$

$$\mathcal{H}_0 = \frac{1}{4} t_0 \left[(2 + x_0) \rho_0^2 - (2x_0 + 1) \sum_q \rho_q^2 \right] = \sum_{T=0,1} C_T [\rho_0] \rho_T^2$$

$$\mathcal{H}_3 = \frac{1}{24} t_3 \rho_0^\alpha \left[(2 + x_3) \rho_0^2 - (2x_3 + 1) \sum_q \rho_q^2 \right]$$

$$\mathcal{H}_{\text{eff}} = \frac{1}{8} [t_1(2 + x_1) + t_2(2 + x_2)] \tau_0 \rho_0 = \sum_{T=0,1} C_T^\tau \tau_T \rho_T$$

$$+ \frac{1}{8} [t_2(2x_2 + 1) - t_1(2x_1 + 1)] \sum_q \tau_q \rho_q$$

$$\mathcal{H}_{\text{fin}} = \frac{1}{32} [t_2(2 + x_2) - 3t_1(2 + x_1)] \rho_0 \Delta \rho_0 = \sum_{T=0,1} C_T^{\Delta \rho} \rho_T \Delta \rho_T$$

$$+ \frac{1}{32} [3t_1(2x_1 + 1) + t_2(2x_2 + 1)] \sum_q \rho_q \Delta \rho_q \quad \mathbf{C_0^{\nabla J} \neq 3C_1^{\nabla J} : SLy10}$$

$$\mathcal{H}_{\text{so}} = -\frac{W_0}{2} \left[\rho_0 \nabla \cdot \mathbf{J}_0 + \sum_q \rho_q \nabla \cdot \mathbf{J}_q \right] = \sum_{T=0,1} C_T^{\nabla J} \rho_T \nabla \cdot \mathbf{J}_T$$

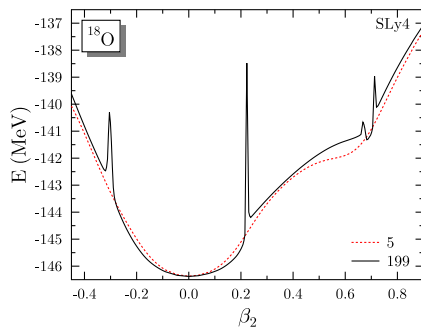
$$\mathcal{H}_{\text{sg}} = -\frac{t_1 x_1 + t_2 x_2}{16} \mathbf{J}_0^2 + \frac{t_1 - t_2}{16} \sum_q \mathbf{J}_q^2 = \sum_{T=0,1} C_T^J \mathbf{J}_T^2 : \mathbf{SLy4}$$

- Different interaction used in the pairing channel

$$V_p(\mathbf{r}) = V_0 \left[1 - \left(\frac{\rho_0(\mathbf{r})}{\rho_{\text{sat}}} \right)^{\alpha'} \right] \delta(\mathbf{r})$$

- *Unconvenient* terms sometimes disregarded: $\mathbf{s}_0 \cdot \Delta \mathbf{s}_0$, $\mathbf{s}_1 \cdot \Delta \mathbf{s}_1$.
- Fit of the parameters on infinite nuclear matter properties on (ground state) properties of (doubly magic) nuclei
 - Good results for ground state properties of even-even nuclei (but how good is good ?)
 - Encouraging constrained HF and beyond mean-field results
- Several annoying technical questions
 - Coupling constants of the time odd part of the functional
 - Calculations sometimes do not converge for some functional
 - How to deal with the density dependent terms in beyond mean-field calculations ?

■ Beyond mean field calculations with a Skyrme EDF



Poles
and steps
in the
projected
energy

See: PRC 79, 044318, 044319 and 044320.

- ⇒ Density dependent term must be dropped...
- ⇒ Three-body interaction ?
- ⇒ Four-body interaction ?
- ⇒ In Hartree, Fock and pairing terms...

New strong constraints on the EDF

The EDF must be derived from an interaction: $E = \langle \hat{T} + \hat{V} \rangle$

- No density dependence
- All terms kept in the functional (Hartree, Fock and pairing)
- Must give attractive pairing

But that's not all:

- First, mean-field calculations must give converged results

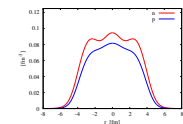
See:

A. Pastore, D. Davesne, K.B., J. Meyer and V. Hellemans,
Phys. Scr. **2013**, 014014.

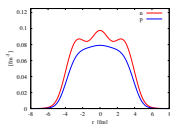
A. Pastore, K.B., D. Davesne, J. Meyer,
Int. J. of Modern Phys. E Vol. 21, No. 05, 1250040.

- Instabilities often experienced with the Skyrme functionals
 - Ferromagnetic instabilities: (spin polarization) $n \uparrow, p \uparrow$
 - Isospin instabilities: neutron-proton *segregation*
 - Both: $n \uparrow, p \downarrow$

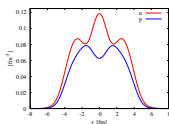
- Example: isospin instability in ^{48}Ca



$$C_1^{\Delta\rho} = 15 \text{ MeV fm}^5 \\ \sim \text{SLy5}$$



$$25 \text{ MeV fm}^5$$



$$35 \text{ MeV fm}^5$$



$$\gtrsim 36 \text{ MeV fm}^5$$

T. Lesinski, K.B., T. Duguet, J. Meyer, PRC 74, 044315 (2006).

Linear response – Instabilities in infinite nuclear matter

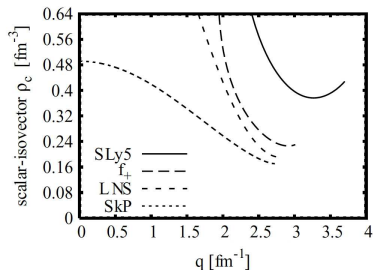
Response of the system to a perturbation given by

$$\mathcal{Q}^{(\alpha)} = \sum_a e^{i\mathbf{q}\cdot\mathbf{r}_a} \Theta_a^{(\alpha)},$$
$$\Theta_a^{\text{SS}} = 1_a, \quad \Theta_a^{\text{VS}} = \boldsymbol{\sigma}_a, \quad \Theta_a^{\text{SV}} = \vec{\tau}_a, \quad \Theta_a^{\text{VV}} = \boldsymbol{\sigma}_a \vec{\tau}_a$$

Response functions are given by

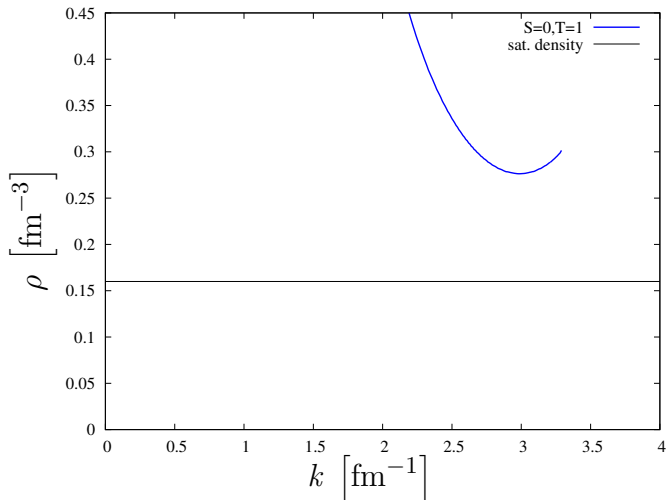
$$\chi^{(\alpha)}(\omega, \mathbf{q}) = \frac{1}{\Omega} \sum_n |\langle n | \mathcal{Q}^{(\alpha)} | 0 \rangle|^2 \left(\frac{1}{\omega - E_{n0} + i\eta} - \frac{1}{\omega + E_{n0} - i\eta} \right)$$

(Cf. C. Garcia-Recio *et al.*, *Ann. of Phys.* 214 (1992) 293–340)



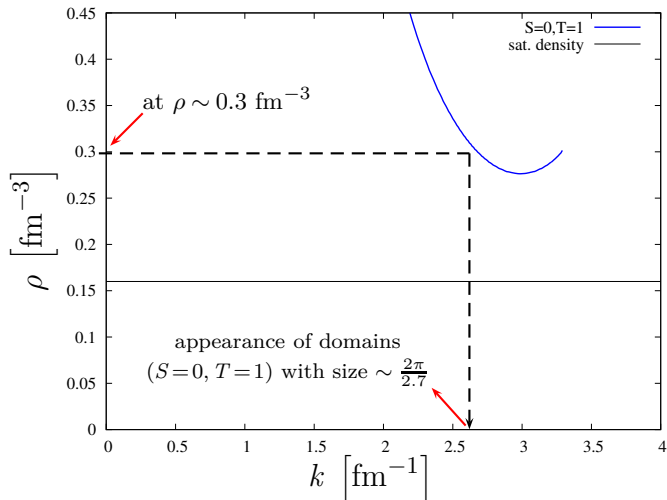
- Predicts instabilities in finite size systems
- Easy to implement
- Negligible computation time
- Might be crucial with tensor, 3- or 4-body terms

Pole of the response at $E = 0 \equiv$ instability



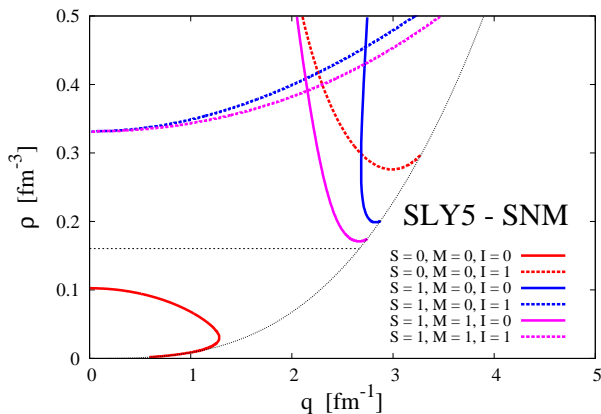
- T. Lesinski, K.B., T. Duguet, J. Meyer, PRC 74, 044315 (2006);
- D. Davesne, M. Martini, K.B., J. Meyer, Phys. Rev. C80, 024314 (2009),
erratum: Phys. Rev. C 84, 059904(E) (2011).

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“Standard” EDFs, instabilities and Murphy’s law



Murphy’s law: “*Anything that can go wrong will go wrong*”.

Instabilities are **very difficult** to detect with a code working on an
harmonic oscillator basis

New Skyrme effective interaction for mean-field and beyond mean-field calculations

- We need
 - An interaction (no density dependence)
 - That can be used in all channels (attractive pairing)
 - Stable in all spin/isospin channel

- Previous work (thesis of J. Sadoudi):
 - Finite size instability can be avoided and correct reproduction of masses can be achieved with 2- and 3-body terms
 - A 4-body might help
 - Properties in the pairing channel not considered at that time.

Skyrme effective interaction with 2-, 3- and 4-body terms

(Cf J. Sadoudi thesis, CEA Saclay)

■ Two-body effective interaction

$$\begin{aligned}V_{\text{eff}} &= t_0 (1 + x_0 \hat{P}^\sigma) \delta(\mathbf{r}) && \text{local} \\ &+ \frac{1}{2} t_1 (1 + x_1 \hat{P}^\sigma) (\mathbf{k}'^2 \delta(\mathbf{r}) + \delta(\mathbf{r}) \mathbf{k}^2) && \text{non local} \\ &+ t_2 (1 + x_2 \hat{P}^\sigma) \mathbf{k}' \cdot \delta(\mathbf{r}) \mathbf{k} && \text{non local} \\ &+ i W_0 \hat{\sigma} \cdot [\mathbf{k}' \times \delta(\mathbf{r}) \mathbf{k}] && \text{spin-orbit}\end{aligned}$$

■ Complemented it with

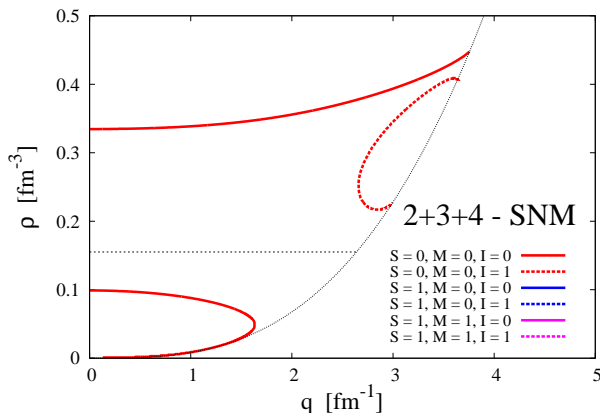
$$\begin{aligned}&3 u_0 \delta(\mathbf{r}_{12}) \delta(\mathbf{r}_{13}) \\ &+ \frac{3}{2} u_1 (1 + y_1 \hat{P}^\sigma) \left[\delta(\mathbf{r}_{12}) \delta(\mathbf{r}_{13}) \mathbf{k}_{12}^2 + \mathbf{k}_{12}'^2 \delta(\mathbf{r}_{12}) \delta(\mathbf{r}_{13}) \right] \\ &+ 3 u_2 (1 + y_{21} \hat{P}_{12}^\sigma + y_{22} \hat{P}_{13}^\sigma) \mathbf{k}_{12}' \cdot \delta(\mathbf{r}_{12}) \delta(\mathbf{r}_{13}) \mathbf{k}_{12} \\ &+ v_0 \delta(\mathbf{r}_{12}) \delta(\mathbf{r}_{13}) \delta(\mathbf{r}_{14})\end{aligned}$$

■ And possibly: tensor, D-wave and 3-body spin-orbit...

Interactions with 2-, 3- and 4-body terms

Stability: How do the 3- and 4-body terms change the picture ?

→ They generate nicer figures !



No hint in the Landau parameters !

Tentative fit with simplified interaction: SLyMR0

See: Phys. Scr. 2013 014013

J. Sadoudi, M. Bender, K.B., D. Davesne, R. Jodon, T. Duguet

- 3- and 4-body terms reduced to

$$3 u_0 \delta(\mathbf{r}_{12})\delta(\mathbf{r}_{13}) + v_0 \delta(\mathbf{r}_{12})\delta(\mathbf{r}_{13})\delta(\mathbf{r}_{14})$$

- Infinite nuclear matter properties

- $\rho_{\text{sat}} = 0.152 \text{ fm}^{-3}$,
- $E/A = -15.04 \text{ MeV}$,
- $K_{\infty} = 264.2 \text{ MeV}$,
- $m^*/m = 0.47$,
- $J = 23 \text{ MeV}$.

- Allows to test the mean-field and beyond mean-field machinery:
Beyond mean-field calculations (B. Bally, M. Bender) in
progress...

SLyMR0: Results

Skyrme EDF for
beyond MF

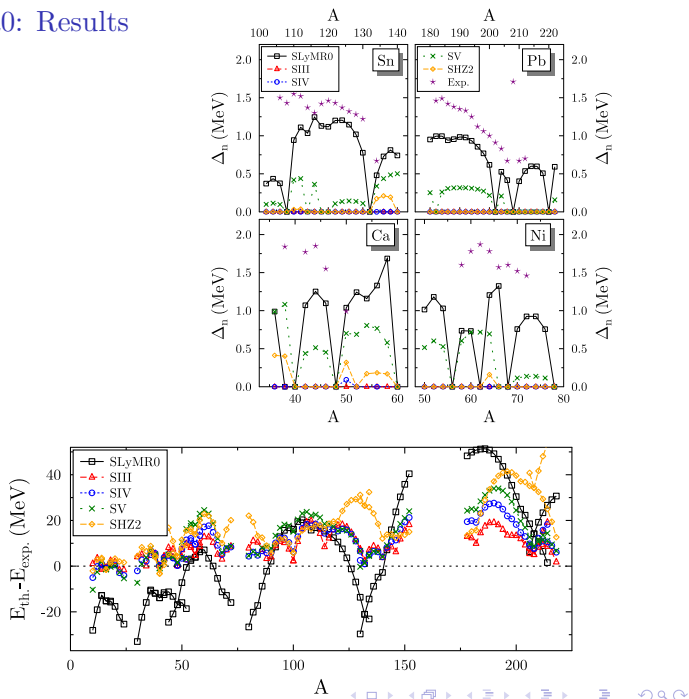
K. Bennaceur

[Introduction](#)

[Skyrme EDFs](#)

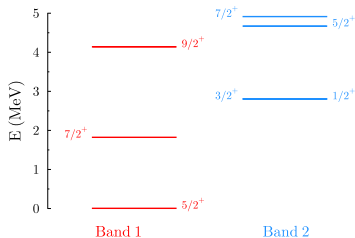
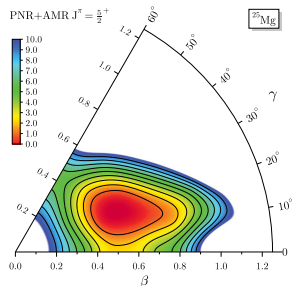
[New constraints](#)

[New Skyrme
interaction](#)



A

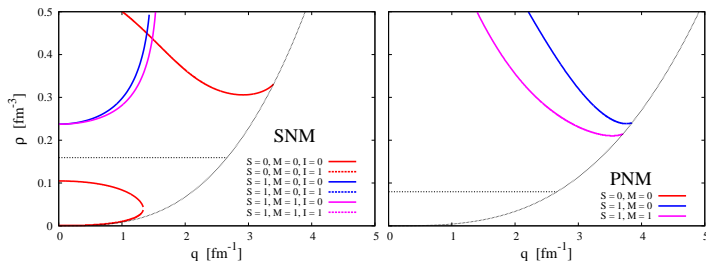
SLyMR0: Results for ^{25}Mg



Calculation for odd-even nucleus with blocking and
projection on N, Z and J

Work in progress: SLyMR1 interaction

- Interaction with 2-, 3- and 4-body terms
- Used **both** in the ph and pp channels
- Perfectly stable !
- Incompressibility too high for the moment



$\rho_{\text{sat}} = 0.16 \text{ fm}^{-3}$	$E/A = -16.1 \text{ MeV}$	$K_{\infty} = 315 \text{ MeV}$	$m^*/m = 0.62$
$a_I = 28.8 \text{ MeV}$	$L = 49.7 \text{ MeV}$	$Q = -376 \text{ MeV}$	$\frac{\Delta m^*}{m} = 0.29 > 0$
$F_0 = -0.12$	$F'_0 = 0.46$	$G_0 = -0.36$	$G'_0 = 0.26$
$F_1 = -1.14$	$F'_1 = 1.37$	$G_1 = -0.63$	$G'_1 = 0.16$

- Attractive pairing but too weak...

Pairing is built from all terms of the interaction, where the attraction comes from ? What can be tuned to enhance it ?

- What about surface properties ?

How the 2- and 3-body gradient terms act on the surface ?

→ Constraint on surface energy will be added

- Four-body term: is it really needed ?

Life is complicated and the four-body term may have driven us to a bad region of parameters

→ Attempt to make a fit without the 4-body term.

- **IPN Lyon:** K. Bennaceur, D. Davesne, R. Jodon, J. Meyer
- **CENBG:** B. Avez, B. Bally, M. Bender, J. Sadoudi
- **IRFU:** T. Duguet
- **ULB:** V. Heelemans, P.H. Heenen, A. Pastore, M. Martini
- **UW:** T. Lesinski